

## Testing the dark origin of neutrino masses with oscillation experiments

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## Dynamic neutrino mass?



Some motivations: Lorenz et al. (PRD, 2021)
➢ Redshift-dependent bounds on ∑m<sub>v</sub> from CMB, BAO and SN
➢ Recent DESI results indicate some tension in ∑m<sub>v</sub> DESI Collaboration (JCAP, 2025)
➢ Relating the two puzzles

For ultralight DM  $10^{-19} \lesssim m_{\phi} \ll 10 \text{eV}$  $\Delta m_{ij}^2 \sim \Delta m_{ijD}^2(\boldsymbol{x}) \cos^2(m_{\phi}t)$ 

DM density- Time dependent modulation assuming relativistic neutrinos For specific realization of "dark" neutrino mass, see: Capozzi, Shoemaker and Vecchi (JCAP, 2018) Choi, Chun and Kim (Phys.Dark Univ., 2020) Huang, Lindner, Martinez-Mirave and Sen (PRD, 2022) ChoeJo, Kim and Lee (PRD, 2023) Sen and Smirnov (JCAP, 2024) Plestid and Tevosyan (2024) Lee (2024)

## Dark matter mass $<< 10^{-14} \text{ eV}$



Time averaged probabilities
 Constant amplitudes



Vacuum mass is favored at  $4.5\sigma$ 

Data from KamLAND experiment

## Dark matter mass >> $10^{-14} \,\mathrm{eV}$ $T_{\mathrm{exp}} \gg t_{\mathrm{dB}}$

Both time- and space-averaged probabilities





Oscillation amplitudes are suppressed in a model-independent way